

**COMPACT SPACE-LIKE HYPERSURFACES
WITH CONSTANT SCALAR CURVATURE
IN LOCALLY SYMMETRIC LORENTZ SPACES**

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ABSTRACT. A new class of $(n + 1)$ -dimensional Lorentz spaces of index 1 is introduced which satisfies some geometric conditions and can be regarded as a generalization of Lorentz space form. Then, the compact space-like hypersurface with constant scalar curvature of this spaces is investigated and a gap theorem for the hypersurface is obtained.

1. INTRODUCTION

Let \mathbb{N}_p^{n+p} be an $(n + p)$ -dimensional connected semi-Riemannian manifold of index p . It is called a semi-definite space of index p . When we refer to index p , we mean that there are only p negative eigenvalues of semi-Riemannian metric of \mathbb{N}_p^{n+p} and the other eigenvalues are positive. In particular, \mathbb{N}_1^{n+1} is called a Lorentz space when $p = 1$. When the Lorentz space \mathbb{N}_1^{n+1} is of constant curvature c , we call it Lorentz space form, denote it by $\mathbb{N}_1^{n+1}(c)$, with de Sitter space $\mathbb{S}_1^{n+1}(1)$ and anti-de Sitter space $\mathbb{H}_1^{n+1}(-1)$ as its special cases. A hypersurface M of a Lorentz space is said to be space-like if the induced metric from that of the ambient space is positive definite.

The authors in [3] introduced a class of Lorentz spaces \overline{M} of index 1. Let $\overline{\nabla}$, \overline{K} and \overline{R} denote the semi-Riemannian connection, sectional curvature and curvature tensor on \overline{M} , respectively. For constant c_1 , c_2 and c_3 , they considered Lorentz spaces which satisfy the following conditions:

- (1) for any space-like vector u and any time-like vector v , $\overline{K}(u, v) = -\frac{c_1}{n}$,
- (2) for any space-like vector u and v , $\overline{K}(u, v) \geq c_2$,
- (3)

$$|\overline{\nabla} \overline{R}| \leq \frac{c_3}{n}.$$

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When \overline{M} satisfies conditions (1) and (2), they say that \overline{M} satisfies condition (*). When \overline{M} satisfies conditions (1) – (3), they say that \overline{M} satisfies condition (**).

Also they give some examples as following.

Example 1.1. The semi-Riemannian product manifold $H_1^k(-\frac{c_1}{n}) \times M^{n+1-k}(c_2)$, $c_1 > 0$. Its sectional curvature is given by

$$\overline{K}(u_1, u_b) = \overline{K}(u_a, u_b) = -\frac{c_1}{n}, \quad \overline{K}(u_a, u_r) = 0, \quad \overline{K}(u_r, u_s) = c_2,$$

where $a, b = 2, \dots, k$; $r, s = k + 1, \dots, n + 1$, u_1 and u_a, u_r denote time-like and space-like vectors respectively.

Example 1.2. The semi-Riemannian product manifold $R_1^k \times S^{n+1-k}(1)$. Its sectional curvature is given by

$$\overline{K}(u_1, u_a) = \overline{K}(u_a, u_b) = 0, \quad \overline{K}(u_1, u_r) = 0, \quad \overline{K}(u_r, u_s) = 1,$$

where $a, b = 2, \dots, k$; $r, s = k + 1, \dots, n + 1$. In particular, $R_1^1 \times S^n(1)$ is called Einstein Static Universe. Notice that it is not a Lorentz space form.

The authors in [2, 8] investigated complete space-like hypersurfaces M in a Lorentz space satisfying condition (**). They estimate the square norm of the second fundamental form of M under some conditions. Baek-Cheng-Suh in [3] studied complete space-like hypersurfaces with constant mean curvature satisfying the condition (*). Later, Xu and Chen in [9] generalized the related results in [3] by investigating complete space-like submanifolds with constant mean curvature in locally symmetric semi-Riemannian spaces. Recently, Liu and Wei in [4] obtained a gap theorem for complete space-like hypersurface with constant scalar curvature in locally symmetric Lorentz spaces.

Now we consider Lorentz spaces which satisfy another condition:

(4) for any space-like vectors u and v , $\overline{K}(u, v) \leq c_2$.

When \overline{M} satisfies conditions (1) and (4), we shall say that \overline{M} satisfies conditions ($\overline{*}$). When \overline{M} satisfies conditions (1), (3) and (4), we shall say that \overline{M} satisfies condition ($\overline{**}$). In this paper, we mainly discuss the compact space-like hypersurfaces with constant scalar curvature in a locally symmetric Lorentz spaces satisfying the condition ($\overline{*}$). It is worthy to point out that both Example 1.1 and 1.2 satisfy the condition ($\overline{*}$).

Remark 1.3. It is easy to see that a Lorentz space form $N_1^{n+1}(s)$ satisfies both conditions (***) and ($\overline{**}$), where $-\frac{c_1}{n} = c_2 = s$.

Remark 1.4. If a Lorentz space \overline{M} is locally symmetric, then the condition (3) holds naturally, because $\overline{\nabla} \overline{R} = 0$ in this situation.

Remark 1.5. As discussed in section 4, our theorem extend the results in [6] under some geometric conditions.

2. PRELIMINARIES

Let $(\overline{M}, \overline{g})$ be an $(n + 1)$ -dimensional Lorentz space of index 1. Throughout the paper, manifolds are assumed to be connected and geometric objects are assumed

to be of class C^∞ . For any point $p \in \overline{M}$, we choose a local field of semi-orthonormal frames $\{e_A\} = \{e_1, e_2, \dots, e_{n+1}\}$ on a neighborhood of p , where e_1, \dots, e_n are space-like and e_{n+1} is time-like. We use the following convention on the range of indices throughout the paper

$$A, B, \dots = 1, \dots, n + 1; \quad i, j, \dots = 1, 2, \dots, n.$$

Let $\{\omega_A\} = \{\omega_1, \omega_2, \dots, \omega_{n+1}\}$ denote the dual frame fields of $\{e_A\}$ on \overline{M} . The metric tensor \overline{g} of \overline{M} satisfies $\overline{g}(e_A, e_B) = \epsilon_A \delta_{AB}$, where $\epsilon_1 = \dots = \epsilon_n = 1$ and $\epsilon_{n+1} = -1$. The canonical forms $\{\omega_A\}$ and the connection forms $\{\omega_{AB}\}$ satisfy the following structure equations

$$(2.1) \quad d\omega_A = - \sum_B \epsilon_B \omega_{AB} \wedge \omega_B, \quad \omega_{AB} + \omega_{BA} = 0,$$

$$(2.2) \quad d\omega_{AB} = - \sum_C \epsilon_C \omega_{AC} \wedge \omega_{CB} - \frac{1}{2} \sum_{C,D} \epsilon_C \epsilon_D \overline{R}_{ABCD} \omega_C \wedge \omega_D.$$

The components \overline{R}_{CD} of the Ricci tensor and the scalar curvature \overline{R} are given respectively by

$$(2.3) \quad \overline{R}_{CD} = \sum_B \epsilon_B \overline{R}_{BCDB},$$

and

$$(2.4) \quad \overline{R} = \sum_A \epsilon_A \overline{R}_{AA}.$$

The components $\overline{R}_{ABCD;E}$ of the covariant derivative of the Riemannian curvature tensor \overline{R} are defined by

$$(2.5) \quad \sum_E \epsilon_E \overline{R}_{ABCD;E} \\ = d\overline{R}_{ABCD} - \sum_E \epsilon_E (\overline{R}_{EBCD} \omega_{EA} + \overline{R}_{AECD} \omega_{EB} + \overline{R}_{ABED} \omega_{EC} + \overline{R}_{ABCE} \omega_{ED}).$$

Restricting the forms $\{\omega_A\}$ to a space-like hypersurface M in \overline{M} , we have

$$(2.6) \quad \omega_{n+1} = 0,$$

and the induced metric g of M is given by $g = \sum_i \omega_i \otimes \omega_i$. It is well known that by Cartan's Lemma we get

$$(2.7) \quad \omega_{(n+1)i} = \sum_j h_{ij} \omega_j, \quad h_{ij} = h_{ji},$$

where h_{ij} are the coefficients of the second fundamental form of M . Then we denote by $H = \frac{1}{n} \sum_i h_{ii}$ and $S = \sum_{ij} h_{ij}^2$ the mean curvature and squared norm of the second fundamental form of M , respectively.

The structure equations of M are given by

$$(2.8) \quad d\omega_i = -\sum_i \omega_{ij} \wedge \omega_j, \quad \omega_{ij} + \omega_{ji} = 0,$$

$$(2.9) \quad d\omega_{ij} = -\sum_k \omega_{ik} \wedge \omega_{kj} - \frac{1}{2} \sum_{k,l} R_{ijkl} \omega_k \wedge \omega_l.$$

The Gauss equation is given by

$$(2.10) \quad R_{ijkl} = \bar{R}_{ijkl} + (h_{ik}h_{jl} - h_{il}h_{jk}).$$

The Ricci tensor and normalized scalar curvature of M are given respectively by

$$(2.11) \quad R_{ij} = \sum_k \bar{R}_{kij} - nHh_{ij} + \sum_k h_{ik}h_{kj},$$

and

$$(2.12) \quad n(n-1)R = \sum_{j,k} \bar{R}_{kjj} - n^2H^2 + S.$$

Let \bar{M} be a locally symmetric Lorentz space satisfying the condition (\ast) . We know that the scalar curvature \bar{R} of \bar{M} is a constant. By using the structure equations of \bar{M} , we have

$$(2.13) \quad \bar{R} = \sum_A \epsilon_A \bar{R}_{AA} = -2 \sum_i \bar{R}_{(n+1)ii(n+1)} + \sum_{i,j} \bar{R}_{ijji} = -2c_1 + \sum_{i,j} \bar{R}_{ijji},$$

which means that $\sum_{i,j} \bar{R}_{ijji}$ is a constant. We assume from now that the scalar curvature R of M is constant. Together with the above equation and (2.12), we define a constant P by

$$(2.14) \quad n(n-1)P = n^2H^2 - S = \sum_{ij} \bar{R}_{ijji} - n(n-1)R.$$

By taking exterior differentiation of (2.7) and defining h_{ijk} by

$$(2.15) \quad \sum_k h_{ijk} \omega_k = dh_{ij} - \sum_k (h_{kj} \omega_{ki} + h_{ik} \omega_{kj}),$$

we have the following Codazzi equation

$$(2.16) \quad h_{ijk} - h_{ikj} = \bar{R}_{(n+1)ijk}.$$

Similarly, we define h_{ijkl} by

$$(2.17) \quad \sum_l h_{ijkl} \omega_l = dh_{ijk} - \sum_l (h_{ljk} \omega_{li} + h_{ilk} \omega_{lj} + h_{ijl} \omega_{lk}).$$

By taking exterior differentiation of (2.15), we have Ricci formula for the second fundamental form of M

$$(2.18) \quad h_{ijkl} - h_{ijlk} = -\sum_r (h_{ir} R_{rjkl} + h_{jr} R_{rikl}).$$

Restricting (2.5) on M , $\bar{R}_{(n+1)ijk;l}$ is given by

$$(2.19) \quad \begin{aligned} \bar{R}_{(n+1)ijk;l} &= \bar{R}_{(n+1)ijkl} + \bar{R}_{(n+1)i(n+1)k} \bar{h}_{jl} \\ &+ \bar{R}_{(n+1)ij(n+1)} h_{kl} + \sum_m \bar{R}_{mijk} h_{ml}, \end{aligned}$$

where $\bar{R}_{(n+1)ijkl}$ denote the covariant derivative of $\bar{R}_{(n+1)ijk}$ as a tensor on M by

$$(2.20) \quad \begin{aligned} \sum_l \bar{R}_{(n+1)ijkl} \omega_l &= d\bar{R}_{(n+1)ijk} - \sum_l \bar{R}_{(n+1)ljk} \omega_{li} \\ &- \sum_l \bar{R}_{(n+1)ilk} \omega_{lj} - \sum_l \bar{R}_{(n+1)ijl} \omega_{lk}. \end{aligned}$$

Remark 2.1. If \bar{M} is a Lorentz space form of index 1, by a straightforward calculation we check that the sum of the last three terms of right-hand side of (2.19) goes to zero. Then we have $\bar{R}_{(n+1)ijk;l} = \bar{R}_{(n+1)ijkl}$, which is the same as in the case that the ambient space is a space form.

It is well known that the Laplacian Δh_{ij} is defined by

$$(2.21) \quad \Delta h_{ij} = \sum_k h_{ijkk}.$$

By using Codazzi equation and Ricci formula, we get

$$(2.22) \quad \begin{aligned} \Delta h_{ij} &= \sum_k h_{ikjk} + \sum_k \bar{R}_{(n+1)ijkk} = \sum_k h_{kij k} + \sum_k \bar{R}_{(n+1)ijkk} \\ &= \sum_k \left(h_{kikj} - \sum_l (h_{kl} R_{lij k} + h_{il} R_{lkj k}) + \bar{R}_{(n+1)ijkk} \right). \end{aligned}$$

From the Codazzi equation $h_{ikjk} = h_{kkij} + \bar{R}_{(n+1)kikj}$, we have

$$\Delta h_{ij} = \sum_k h_{kkij} + \sum_k (\bar{R}_{(n+1)ijkk} + \bar{R}_{(n+1)kikj}) - \sum_{k,l} (h_{kl} R_{lij k} + h_{il} R_{lkj k}).$$

Together with Gauss equation and above equation and (2.19), we have

$$(2.23) \quad \begin{aligned} \Delta h_{ij} &= \sum_k h_{kkij} + \sum_k (\bar{R}_{(n+1)ijk;k} + \bar{R}_{(n+1)kik;j}) \\ &- \sum_{k,l} (2h_{kl} \bar{R}_{lij k} + h_{jl} \bar{R}_{lkik} + h_{il} \bar{R}_{lkjk}) + S h_{ij} \\ &- \sum_k (h_{kk} \bar{R}_{(n+1)ij(n+1)} + h_{ij} \bar{R}_{(n+1)k(n+1)k}) - nH \sum_l h_{il} h_{jl}. \end{aligned}$$

Thus

$$\begin{aligned}
 \frac{1}{2} \Delta S &= \sum_{i,j,k} h_{ijk}^2 + \sum_{i,j} h_{ij} \Delta h_{ij} \\
 &= \sum_{i,j,k} h_{ijk}^2 + \sum_{i,j} (nH)_{ij} h_{ij} + \sum_{i,j,k} (\bar{R}_{(n+1)ijk;k} + \bar{R}_{(n+1)kik;j}) h_{ij} \\
 (2.24) \quad &\quad - \left(nH \sum_{i,j} h_{ij} \bar{R}_{(n+1)ij(n+1)} + S \sum_k \bar{R}_{(n+1)k(n+1)k} \right) + S^2 \\
 &\quad - \sum_{i,j,k,l} 2(h_{kl} h_{ij} \bar{R}_{lijk} + h_{il} h_{ij} \bar{R}_{lkjk}) - nH \sum_{i,j,l} h_{il} h_{lj} h_{ij}.
 \end{aligned}$$

3. ESTIMATES OF LAPLACIAN AND KEY LEMMAS

Let \bar{M} be a locally symmetric Lorentz space, i.e., $\bar{R}_{ABCD;E} = 0$. We also may choose a canonical bases $\{e_1, e_2, \dots, e_n\}$ such that $h_{ij} = \lambda_i \delta_{ij}$, thus

$$(3.1) \quad \bar{R}_{(n+1)ijk;k} + \bar{R}_{(n+1)kik;j} = 0.$$

Noticing that \bar{M} satisfies condition $(\bar{*})$, we have

$$\begin{aligned}
 &\quad - \left(nH \sum_{i,j} h_{ij} \bar{R}_{(n+1)ij(n+1)} + S \sum_k \bar{R}_{(n+1)k(n+1)k} \right) \\
 (3.2) \quad &= - \left(nH \sum_i \lambda_i \bar{R}_{(n+1)ii(n+1)} + S \sum_i \bar{R}_{(n+1)i(n+1)i} \right) \\
 &= c_1(S - nH^2).
 \end{aligned}$$

Also we have

$$\begin{aligned}
 &\quad - \sum_{i,j,k,l} 2(h_{kl} h_{ij} \bar{R}_{lijk} + h_{il} h_{ij} \bar{R}_{lkjk}) \\
 (3.3) \quad &= - 2 \sum_{j,k} (\lambda_j \lambda_k - \lambda_k^2) \bar{R}_{kjjk} \leq c_2 \sum_{j,k} (\lambda_j - \lambda_k)^2 = 2c_2(nS - n^2H^2).
 \end{aligned}$$

Substituting (3.1), (3.2) and (3.3) in to (2.24), it yields that

$$(3.4) \quad \frac{1}{2} \Delta S \leq \sum_{i,k} h_{iik}^2 + \sum_i \lambda_i (nH)_{ii} + (2nc_2 + c_1)(S - nH^2) + (S^2 - nH \sum_i \lambda_i^3).$$

Lemma 3.1 ([7]). *Let $\{\mu_1, \mu_2, \dots, \mu_n\}$ be real numbers satisfying $\sum_i \mu_i = 0$ and $\sum_i \mu_i^2 = A$, where A is a constant no less than zero. Then we have*

$$\left| \sum_i \mu_i^3 \right| \leq \frac{n-2}{\sqrt{n(n-1)}} A^{\frac{3}{2}},$$

and the equality holds if and only if at least $n-1$ of the μ_i are equal, i.e.,

$$\mu_1 = \mu_2 = \dots = \mu_{n-1} = -\sqrt{\frac{1}{n(n-1)}} A, \quad \mu_n = \sqrt{\frac{n-1}{n}} A.$$

Lemma 3.2. *Let M be a space-like hypersurface with constant normalized scalar curvature R in locally symmetric $(n + 1)$ -dimensional Lorentz space satisfying the condition $(\bar{*})$. If $h_{ijk} \geq 0$, then*

$$\sum_{i,j,k} h_{ijk}^2 \leq n^2 |\nabla H|^2.$$

Proof. Notice that the following equation holds:

$$\begin{aligned} n^2 |\nabla H|^2 &= \sum_k \left(\sum_{i,j} h_{ijk} \right)^2 = \sum_{i,j,k,l,m} h_{ijk} h_{lmk} \\ &= \sum_{i,j,k} h_{ijk}^2 + \sum_{i \neq l,j,k,m} h_{ijk} h_{lmk} + \sum_{i,j \neq m,k} h_{ijk} h_{imk}. \end{aligned}$$

Then the proof follows from the above equation. □

Next we will use the well known self-adjoint operator \square introduced in [1] to the function nH and using (2.14), we have

$$\begin{aligned} \square(nH) &:= \sum_{i,j} (nH \delta_{ij} - h_{ij})(nH)_{ij} \\ (3.5) \quad &= \frac{1}{2} \Delta(nH)^2 - \sum_i (nH)_i^2 - \sum_i \lambda_i (nH)_{ii} \\ &= \frac{1}{2} \Delta(n(n-1)P) + \frac{1}{2} \Delta S - n^2 |\nabla H|^2 - \sum_i \lambda_i (nH)_{ii}. \end{aligned}$$

By (2.14), we know that P is a constant, so we have $\frac{1}{2} \Delta(n(n-1)P) = 0$. Then substituting (3.4) to (3.5), we obtain

$$(3.6) \quad \square(nH) \leq \sum_{i,j,k} h_{ijk}^2 - n^2 |\nabla H|^2 + (2nc_2 + c_1)(S - nH^2) + (S^2 - nH \sum_i \lambda_i^3).$$

Lemma 3.3. *Let M be a compact space-like hypersurface of dimension n with constant scalar curvature in a locally symmetric Lorentz space which satisfies condition $(\bar{*})$ and $h_{ijk} \geq 0$. Then we have the following inequality*

$$\square(nH) \leq \frac{n-1}{n} (S - nP) \phi_P(S),$$

where $\phi_P(S) = nc - 2(n-1)P + \frac{n-2}{n}S + \frac{n-2}{n} \sqrt{(n(n-1)P + S)(s - nP)}$ and $c = 2c_2 + \frac{c_1}{n}$.

Proof. We denote

$$\mu_i = \lambda_i - H, \quad B = \sum_i \mu_i^2.$$

It is obvious to see that

$$\sum_i \mu_i = 0, \quad B = S - nH^2, \quad \sum_i \lambda_i^3 = \sum_i \mu_i^3 + 3HB + nH^3.$$

By using Lemma 3.1, we have

$$(3.7) \quad \begin{aligned} -nH \sum_i \lambda_i^3 &= -n^2 H^4 - 3nH^2 B - nH \sum_i \mu_i^3 \\ &\leq 2n^2 H^4 - 3nSH^2 + \frac{n(n-2)}{\sqrt{n(n-1)}} \|H\| B^{\frac{3}{2}}. \end{aligned}$$

Substituting (3.7) to (3.6) and together with the Lemma 3.2, we get

$$(3.8) \quad \square(nH) \leq B \left(nc - nH^2 + B + \frac{n(n-2)}{\sqrt{n(n-1)}} \|H\| B^{\frac{1}{2}} \right).$$

It follows from (2.14) that

$$(3.9) \quad B = S - nH^2 = \frac{n-1}{n} (S - nP).$$

Putting the above equation into (3.8), we get

$$(3.10) \quad \square(nH) \leq \frac{n-1}{n} (S - nP) \phi_H(S),$$

where

$$(3.11) \quad \phi_H(S) = nc - 2nH^2 + S + \frac{n(n-2)}{\sqrt{n(n-1)}} \|H\| \sqrt{S - nH^2}.$$

Putting (3.9) into (3.11), we have

$$\phi_P(S) = nc - 2(n-1)P + \frac{n-2}{n} S + \frac{n-2}{n} \sqrt{(n(n-1)P + S)(S - nP)}.$$

Finally, (3.10) becomes

$$(3.12) \quad \square(nH) \leq \frac{n-1}{n} (S - nP) \phi_P(S),$$

then we complete the proof. \square

4. MAIN THEOREMS AND PROOFS

Theorem 4.1. *Let M be a compact space-like hypersurface of dimension n (where $n > 2$) with constant scalar curvature in a locally symmetric Lorentz space of dimension $n + 1$ which satisfies condition $(\bar{*})$ and $h_{ijk} \geq 0$. If $0 \leq c \leq P$ or $c \leq P < \frac{2}{n}c$ or $P > \frac{1}{n-1}c, c < 0$, then the norm square of the second fundamental form S satisfies*

$$S \geq \frac{n}{(n-2)(nP - 2c)} (n(n-1)P^2 - 4c(n-1)P + nc^2),$$

where P is given by (2.14) and $c = 2c_2 + \frac{c_1}{n}$.

Proof. Since \square is a self-adjoint operator and M is compact, then we have

$$(4.1) \quad \int_M \square(nH) * 1 = 0.$$

We notice that $S - nP \geq 0$ holds naturally by (3.9) because $S \geq nH^2$. By taking integration on both sides of (3.12), we get $\phi_P(S) \geq 0$. By directly calculation we see that $\phi_P(S) \geq 0$ is equivalent to

$$S \geq \frac{n}{n-2}(2(n-1)P - nc)$$

or

$$\frac{n}{(n-2)(nP-2c)}(n(n-1)P^2 - 4c(n-1)P + nc^2) \leq S < \frac{n}{n-2}(2(n-1)P - nc).$$

By solving the above inequalities, we complete the proof. □

Theorem 4.2. *Let M be a compact space-like hypersurface of dimension n (where $n > 2$) with constant scalar curvature in a locally symmetric Lorentz space of dimension $n + 1$ which satisfies condition $(\bar{*})$ and $h_{ijk} \geq 0$ and $0 \leq c \leq P$ or $c \leq P < \frac{2}{n}c$ or $P > \frac{1}{n-1}c, c < 0$. If the norm square of the second fundamental form S satisfies*

$$(4.2) \quad nP \leq S \leq \frac{n}{(n-2)(nP-2c)}(n(n-1)P^2 - 4c(n-1)P + nc^2),$$

then

- (i) $S = nP$ and M is totally umbilical, or
- (ii) $S = \frac{n}{(n-2)(nP-2c)}(n(n-1)P^2 - 4c(n-1)P + nc^2)$ and M has two distinct principal curvatures.

Proof. Together with (4.2) and the definition of P , we see that the right-hand side term of (3.12) is non-positive. As in proof of Theorem 4.1, we take integration on both sides of (3.12) and notice (4.1), we have $(S - nP)\phi_P(S) = 0$. In particular, we notice that $\phi_P(S) = 0$ if and only if the equality holds in Lemma 3.1, thus we prove the theorem. □

Remark 4.3. Let \bar{M} in Theorem 4.2 be a Lorentz space form with constant sectional curvature s . In particular, we assume that $s = 1$ such that \bar{M} is nothing but a de Sitter space. As seen in Remark 1.3, we have $-\frac{c_1}{n} = c_2 = 1$. Thus c defined in Lemma 3.3 is 1. Then our theorem is just like Liu’s corollary in [6].

Finally, we discuss the compact space-like surface in a locally symmetric Lorentz spaces of dimension 3, i.e., the version of $n = 2$ of the Theorem 4.1. We using the convention of the ranges of the indexes as following

$$i, j, k = 1, 2, \quad A, B, C = 1, 2, 3.$$

Theorem 4.4. *Let M be a compact space-like surface with constant scalar curvature in a locally symmetric 3-Minkowski space which satisfies condition $(\bar{*})$ and $h_{ijk} \geq 0$. Then*

$$P \leq c,$$

where P is given by (2.14) and $c = 2c_2 + \frac{c_1}{n}$ and h_{ijk} is defined by (2.15).

Proof. We notice that when $n = 2$, (3.12) becomes $\square(2H) \leq (S - 2P)(c - P)$. Taking integration on both sides of the inequality, then we complete the proof. □

Corollary 4.5. *Let M be a compact space-like surface with constant scalar curvature in a locally symmetric 3-Minkowski space which satisfies condition $(\bar{*})$ and $h_{ijk} \geq 0$. If $P \geq c$, then*

- (i) $S = 2P$ and M is totally umbilical, or
- (ii) $P = c$.

The proof is the same as the proof of Theorem 4.2.

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