

Results on quasi-*Einstein metric

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Abstract. We study the quasi-*Einstein metric on Sasakian and (κ, μ) -manifolds. We show that on Sasakian manifolds the *-Ricci operator commutes with tensor field ϕ and quasi-*Einstein Sasakian metric is *-flat. Further, we study $(\kappa < 1, \mu)$ -manifolds with quasi-*Einstein metric and obtain that such manifold is *-flat or locally isometric to $E^{n+1} \times S^n(4)$ or *-Einstein.

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1 Introduction

A Riemannian manifold (M, g) is called *Einstein with $S^* = \nu g$, where S^* denotes *-Ricci tensor and ν is a constant. A *-Ricci soliton is a generalization of *Einstein metric which is given by [13]

$$(1.1) \quad \frac{1}{2} \mathcal{L}_U g + S^* = \nu g,$$

where ν is a constant and U is the potential field. If for a smooth function F , $U = \nabla F$, then (1.1) is called gradient *-Ricci soliton.

The *-Ricci tensor on an almost contact metric manifold M is defined as follows [12]:

$$(1.2) \quad S^*(Z, V) = \frac{1}{2} \text{trace}(X \mapsto R(Z, \phi V)\phi X), \quad \forall Z, V, X \in TM,$$

where ϕ is a $(1, 1)$ -tensor field and R is a Riemann curvature tensor.

The Einstein condition $S = \nu g$ and its generalizations have been studied extensively in contact geometry (see [8, 11, 16]). A generalization of Einstein metric emerged from the m-Bakry-Emery Ricci tensor S_m^F defined as:

$$(1.3) \quad S_m^F = S + \nabla^2 F - \frac{1}{m} dF \otimes dF,$$

has been studied in [4, 5], where $0 < m \leq \infty$, S is the Ricci tensor and $\nabla^2 F$ denotes the Hessian form of F .

A Riemannian manifold (M, g, F, m) is called m -quasi-Einstein [7] if it satisfies

$$(1.4) \quad S + HessF - \frac{1}{m}dF \otimes dF = \nu g,$$

for some $m \in \mathbb{Z}^+$. Similarly, we call (M, g, F, m) , m -quasi-*Einstein if it satisfies

$$(1.5) \quad S^* + HessF - \frac{1}{m}dF \otimes dF = \nu g.$$

If $m = \infty$, then (1.5) gives the gradient *-Ricci soliton. A quasi-*Einstein metric is *Einstein if $F = \text{constant}$. We call a quasi-*Einstein metric steady, expanding, or shrinking, respectively, when $\nu = 0, < 0$ or > 0 .

Sharma [15] proved that a complete K -contact metric with gradient Ricci soliton is a compact Einstein Sasakian manifold and gradient soliton is expanding. As every Sasakian manifold is a K -contact manifold, this result is also true for Sasakian manifolds. Ghosh et al. extended this result for (κ, μ) -spaces [11]. Quasi-Einstein metrics have been studied in extent for a general manifold, and gap results and rigid properties were obtained in (cf. [4, 17]). Further, Ghosh [9] studied quasi-Einstein contact metric on Sasakian manifolds and on (κ, μ) -spaces. Recently, Chen [6] studied the quasi-Einstein metric on almost cosymplectic manifolds.

However, only very little literature is available on the study of *Einstein and its generalization. This inspired us to study quasi-*Einstein structure associated with contact metric manifolds.

2 Preliminaries and some basic results

An odd-dimensional Riemannian manifold M^{2n+1} is called an almost contact metric manifold if it admits a $(1, 1)$ tensor field ϕ , a vector field ξ , a 1-form η and a Riemannian metric g satisfying [1]

$$(2.1) \quad \phi^2 Z = -Z + \eta(Z)\xi, \quad \eta(\xi) = 1,$$

$$(2.2) \quad \eta \circ \phi = 0, \quad \phi\xi = 0,$$

$$(2.3) \quad g(\phi Z, \phi V) = g(Z, V) - \eta(Z)\eta(V), \quad \eta(Z) = g(Z, \xi),$$

where $Z, V \in TM$. A contact manifold is a Riemannian manifold M^{2n+1} with a global 1-form η called a contact 1-form such that $\eta \wedge (d\eta)^n \neq 0$ everywhere on M and $d\eta(Z, V) = g(Z, \phi V)$.

A contact metric manifold is called Sasakian [1] if

$$(\nabla_Z \phi)V = g(Z, V)\xi - \eta(V)Z,$$

where $Z, V \in TM$.

The Riemann curvature tensor R on a Sasakian manifold is given by [1]

$$(2.4) \quad R(Z, V)\xi = \eta(V)Z - \eta(Z)V.$$

On a K -contact manifold [1], we have

$$(2.5) \quad \nabla_Z \xi = -\phi Z,$$

$$(2.6) \quad Q\xi = 2n\xi,$$

$$(2.7) \quad R(Z, \xi)\xi = Z - \eta(Z)\xi.$$

Now, we define self-adjoint operators $h = \frac{1}{2}\mathcal{L}_\xi\phi$ and $l = -R(\xi, \cdot)\xi$, which satisfy [1]:

$$(2.8) \quad \begin{cases} l\xi = 0 = h\xi, & \phi h = -h\phi, & g(hX, Y) = g(hY, X), \\ trh = trh\phi = 0. \end{cases}$$

Also, for any contact metric manifold [1], we have

$$(2.9) \quad \nabla_Z \xi = -\phi Z - \phi hZ,$$

$$(2.10) \quad \nabla_\xi h = -\phi h^2 - \phi l + \phi.$$

A contact metric manifold M^{2n+1} is said to be (κ, μ) -space if Riemann curvature tensor R satisfies [2]

$$(2.11) \quad R(Z, V)\xi = \kappa\{\eta(V)Z - \eta(Z)V\} + \mu\{\eta(V)hZ - \eta(Z)hV\},$$

for some $(\kappa, \mu) \in R^2$. If $\kappa = 1$ and $h = 0$, then (κ, μ) -spaces reduce to the Sasakian manifolds. The non-Sasakian manifolds have proven to be more interesting as there exists the unit tangent sphere bundle of a flat Riemannian manifold with the usual contact metric structure as an example of non-Sasakian spaces of this type. Moreover, this type of space is invariant under D -homothetic transformations. These are the driving factor for the study of this type of manifold. Boeckx proved that non-Sasakian contact metric manifold satisfying (2.11) is completely determined locally by its dimension for the constant values of κ, μ [3].

3 Quasi-*Einstein metric on a Sasakian manifold

In this section, we study the quasi-*Einstein metric on a Sasakian manifold. On a Sasakian manifold M^{2n+1} [10]

$$(3.1) \quad Q^*Z = QZ - (2n - 1)Z - \eta(Z)\xi,$$

$$(3.2) \quad r^* = r - 4n^2,$$

where Q^* , Q , r^* and r are *-Ricci operator, Ricci operator, *-scalar curvature and scalar curvature respectively

Lemma 3.1. [14] *The curvature tensor R of a Riemannian manifold M^{2n+1} with quasi-*Einstein metric satisfies*

$$(3.3) \quad \begin{aligned} R(Z, V)\nabla F &= (\nabla_V Q^*)Z - (\nabla_Z Q^*)V - \frac{\nu}{m}\{(ZF)V - (VF)Z\} \\ &+ \frac{1}{m}\{(ZF)Q^*V - (VF)Q^*Z\}, \end{aligned}$$

$$(3.4) \quad S(V, \nabla F) = \frac{1}{2}Vr^* + \frac{2n\nu}{m}(VF) + \frac{1}{m}\{(Q^*V)F - r^*(VF)\},$$

for any $Z, V \in TM$, where ∇F is gradient of F .

Proof. From (1.5), we have

$$(3.5) \quad Q^*V + \nabla_V \nabla F - \frac{1}{m}(VF)\nabla F = \nu V.$$

We know that

$$(3.6) \quad R(Z, V)X = \nabla_Z \nabla_V X - \nabla_V \nabla_Z X - \nabla_{[Z, V]}X.$$

Using (3.5) in (3.6), we obtain (3.3). Further, contracting (3.3) over Z , we get (3.4). \square

Lemma 3.2. [9] *If $F \in C^\infty$ on a contact metric manifold M such that $dF = (\xi F)\eta$, where d denotes the exterior differentiation. Then F is constant on M .*

Next,

Lemma 3.3. *On a Sasakian manifold M^{2n+1}*

$$(3.7) \quad Q^*\phi = \phi Q^*.$$

Proof. Putting $Z = \xi$ in (3.1) and using (2.6), we obtain

$$(3.8) \quad Q^*\xi = 0.$$

Differentiating (3.8) with respect to Z , we get

$$(3.9) \quad (\nabla_Z Q^*)\xi = Q^*\phi Z.$$

If $V, Z \in \{\xi^\perp\}$, then with respect to ϕ basis $\{Z_i, Z_{n+i} = \phi Z_i, \xi\}$, we have

$$\begin{aligned} g(Q^*\phi Z, \phi V) &= \frac{1}{2} \sum_{i=1}^{2n} g(R(\phi Z, \phi^2 V)\phi Z_i, Z_i) + \frac{1}{2} g(R(\phi Z, \phi^2 V)\phi \xi, \xi) \\ &= \frac{1}{2} \sum_{i=1}^{2n} g(R(V, \phi Z)\phi Z_i, Z_i), \end{aligned}$$

which gives

$$(3.10) \quad g(Q^*\phi Z, \phi V) = g(Q^*V, Z).$$

From (3.1) and using the fact that Q is symmetric, we get

$$(3.11) \quad g(Q^*V, Z) = g(Q^*Z, V).$$

From (3.10) and (3.11), we get

$$-\phi Q^*\phi Z = Q^*Z,$$

wherein multiplying with ' ϕ ' on both sides and using (2.1), we find

$$Q^*\phi Z = \phi Q^*Z,$$

whereby proof is complete. \square

Theorem 3.4. *Let M^{2n+1} be a Sasakian manifold, then $\mathcal{L}_\xi S^* = 0$ and $\nabla_\xi Q^* = 0$.*

Proof. It is well-known that $\mathcal{L}_\xi g = 0$ on a Sasakian manifold. Also, we have

$$(3.12) \quad \mathcal{L}_\xi(R_{Z,\phi V}\phi X) = \mathcal{L}_\xi(\nabla_Z\nabla_{\phi V}\phi X - \nabla_{\phi V}\nabla_Z\phi X - \nabla_{[Z,\phi V]}X),$$

$$(3.13) \quad R_{\mathcal{L}_\xi Z,\phi V}\phi X = \nabla_{\mathcal{L}_\xi Z}\nabla_{\phi V}\phi X - \nabla_{\phi V}\nabla_{\mathcal{L}_\xi Z}\phi X - \nabla_{[\mathcal{L}_\xi Z,\phi V]}\phi X,$$

$$(3.14) \quad R_{Z,\mathcal{L}_\xi\phi V}\phi X = \nabla_Z\nabla_{\mathcal{L}_\xi\phi V}\phi X - \nabla_{\mathcal{L}_\xi\phi V}\nabla_Z\phi X - \nabla_{[Z,\mathcal{L}_\xi\phi V]}\phi X,$$

$$(3.15) \quad R_{Z,\phi V}\mathcal{L}_\xi\phi X = \nabla_Z\nabla_{\phi V}(\mathcal{L}_\xi\phi X) - \nabla_{\phi V}\nabla_Z(\mathcal{L}_\xi\phi X) - \nabla_{[Z,\phi V]}\mathcal{L}_\xi\phi X,$$

$\forall Z, V, X \in TM$.

From (3.12)~(3.15), we obtain

$$\mathcal{L}_\xi(R_{Z,\phi V}\phi X) = R_{\mathcal{L}_\xi Z,\phi V}\phi X + R_{Z,\mathcal{L}_\xi\phi V}\phi X + R_{Z,\phi V}\mathcal{L}_\xi\phi X,$$

which gives $(\mathcal{L}_\xi R)_{Z,\phi V}\phi X = 0$ and contracting it over X , we get $\mathcal{L}_\xi S^* = 0$. This implies

$$(3.16) \quad \mathcal{L}_\xi(g(Q^*Z, V)) - S^*(\mathcal{L}_\xi Z, V) - S^*(Z, \mathcal{L}_\xi V) = 0.$$

Simplifying (3.16), we find

$$(3.17) \quad [\xi, Q^*Z] - Q^*([\xi, Z]) = 0.$$

Using (2.5) in (3.17), we get

$$(3.18) \quad \nabla_\xi Q^* = Q^*\phi - \phi Q^*.$$

Using Lemma 3.3 in (3.18), we obtain the result. \square

Theorem 3.5. *Let M^{2n+1} be a Sasakian manifold satisfying (1.5), then F is constant and quasi-*Einstein soliton is steady.*

Proof. Taking inner product of (3.3) with ξ and using (3.8) and (3.9), we get

$$(3.19) \quad g(R(Z, V)\nabla F, \xi) = -2g(Q^*\phi Z, V) + \frac{\nu}{m}\{(VF)\eta(Z) - (ZF)\eta(V)\}.$$

Putting $V = \xi$ in (3.19) and using (2.4), we get

$$(3.20) \quad \left(\frac{\nu}{m} - 1\right)((\xi F)\eta(Z) - (ZF)) = 0.$$

Now, we claim that $\nu \neq m$. Infact, if $\nu = m$, then putting $Z = \xi$ in (3.3), we get

$$(3.21) \quad \begin{aligned} R(\xi, V)\nabla F &= (\nabla_V Q^*)\xi - (\nabla_\xi Q^*)V - \{(\xi F)V - (VF)\xi\} \\ &+ \frac{1}{m}\{(\xi F)Q^*V - (VF)Q^*\xi\}. \end{aligned}$$

From (2.4), we obtain

$$(3.22) \quad R(\xi, V)\nabla F = (VF)\xi - (\xi F)V.$$

Using (3.22) in (3.21), we get

$$(3.23) \quad (\nabla_V Q^*)\xi - (\nabla_\xi Q^*)V + \frac{1}{m}\{(\xi F)Q^*V - (VF)Q^*\xi\} = 0.$$

Using Theorem 3.4, (3.8) and (3.9) in (3.23), we obtain

$$(3.24) \quad Q^*\phi V + \frac{\xi F}{m}Q^*V = 0.$$

Taking inner product of (3.24) with Z , we find

$$(3.25) \quad g(Q^*\phi V, Z) + \frac{\xi F}{m}g(Q^*V, Z) = 0.$$

Interchanging V and Z in (3.25), we get

$$(3.26) \quad g(Q^*\phi Z, V) + \frac{\xi F}{m}g(Q^*Z, V) = 0.$$

Subtracting (3.26) from (3.25), we obtain

$$(3.27) \quad (Q^*\phi + \phi Q^*)V = 0,$$

$\forall V \in TM$. Using Lemma 3.3 in (3.27), we get $\phi Q^*V = 0$. Which further using (2.1) gives $Q^*V = 0$. Hence $r^* = 0$. Using this in (3.4), we find

$$(3.28) \quad S(V, \nabla F) = 2n(VF).$$

On the other hand from (3.1), we get

$$(3.29) \quad QV = (2n - 1)V + \eta(V)\xi.$$

Using (3.29) in (3.28), we obtain $VF = (\xi F)\eta(V)$. Which gives F constant by use of Lemma 3.2. Therefore, from (1.5), we obtain $\nu = 0$, that gives $m = 0$, a contradiction of the fact that $m > 0$.

Hence $\nu \neq m$ and from (3.20), we have $ZF = (\xi F)\eta(Z)$. By using Lemma 3.2 we find that

$$(3.30) \quad F = \text{constant}.$$

From (3.8), we see that

$$(3.31) \quad S^*(\xi, \xi) = 0.$$

Using (3.30) and (3.31) in (1.5), we get $\nu = 0$. Hence quasi-*Einstein soliton is steady. \square

Corollary 3.6. *Let M^{2n+1} be a Sasakian manifold satisfying (1.5), then M is *-Ricci flat, η -Einstein and scalar curvature is constant.*

Proof. Using $\nu = 0$ and (3.30) in (1.5), we see that

$$(3.32) \quad Q^* = 0.$$

Using (3.32) in (3.1), we find

$$(3.33) \quad QZ = (2n - 1)Z + \eta(Z)\xi.$$

Hence M is η -Einstein. Further, from (3.33) we obtain $r = 4n^2$. Hence scalar curvature is constant. \square

Now, we give an example

Example 3.1. Consider the manifold $M^3 = \{(x, y, z) \in \mathbb{R}^3 : x, y \neq 0\}$ endowed with the structure $\{\phi, \xi, \eta, g\}$

$$(3.34) \quad \begin{cases} \phi(e_1) = e_2, \phi(e_2) = -e_1, \phi(e_3) = 0, \eta = \frac{4}{4+3x^2+3y^2}(ydx - xdy) + dz, \\ e_3 = \xi = \frac{\partial}{\partial z}, g = \frac{1}{(1+\frac{3x^2}{4}+\frac{3y^2}{4})^2}(dx \otimes dx + dy \otimes dy) + \eta \otimes \eta, \\ e_1 = (1 + \frac{3x^2+3y^2}{4})\frac{\partial}{\partial x} - y\frac{\partial}{\partial z}, e_2 = (1 + \frac{3x^2+3y^2}{4})\frac{\partial}{\partial y} + x\frac{\partial}{\partial z}. \end{cases}$$

We subsequently have

$$(3.35) \quad [e_l, e_3] = 0 \text{ for } l = 1, 2; [e_1, e_2] = -\frac{3y}{2}e_1 + \frac{3x}{2}e_2 + 2e_3.$$

$$(3.36) \quad \begin{cases} \nabla_{e_1}e_1 = \frac{3y}{2}e_2, \nabla_{e_2}e_1 = -\frac{3x}{2}e_2 - e_3, \nabla_{e_3}e_1 = -e_2, \nabla_{e_1}e_2 = -\frac{3y}{2}e_1 + e_3, \\ \nabla_{e_2}e_2 = \frac{3x}{2}e_1, \nabla_{e_3}e_2 = e_1, \nabla_{e_1}e_3 = -e_2, \nabla_{e_2}e_3 = e_1, \nabla_{e_3}e_3 = 0. \end{cases}$$

$$(3.37) \quad \begin{cases} R(e_l, e_p)e_s = 0, l \neq p \neq s, l, p, s = 1, 2, 3, R(e_l, e_3)e_3 = e_l, l = 1, 2, \\ R(e_l, e_3)e_l = -e_3, l = 1, 2, R(e_l, e_p)e_p = 0, l \neq p, l, p = 1, 2, \\ S^*(e_l, e_p) = 0, l, p = 1, 2, 3. \end{cases}$$

Further, by using (3.37) in (1.5), we infer the following system of differential equations:

$$(3.38) \quad \begin{cases} \alpha^2 F_{xx} - 2y\alpha F_{zx} + y^2 F_{zz} - \frac{3y\alpha}{2}F_y - \frac{3yx}{2}F_z - \frac{\alpha^2}{m}F_x^2 + \frac{2y\alpha}{m}F_x F_z \\ + \frac{3x\alpha}{2}F_x - \frac{y^2}{m}F_z^2 = \nu; \quad F_{zz} - \frac{1}{m}F_z^2 = \nu, \\ \alpha^2 F_{yy} + 2x\alpha F_{zy} + x^2 F_{zz} - \frac{3x\alpha}{2}F_x + \frac{3xy}{2}F_z + \frac{3y\alpha}{2}F_y - \frac{\alpha^2}{m}F_y^2 \\ - \frac{2x\alpha}{m}F_y F_z - \frac{x^2}{m}F_z^2 = \nu, \\ \frac{3x\alpha}{2}F_y + \alpha^2 F_{xy} - y\alpha F_{zy} + \alpha F_z + x\alpha F_{xz} - yx F_{zz} + \frac{3y\alpha}{2}F_x - \frac{3y^2}{2}F_z - F_z \\ - \frac{\alpha^2}{m}F_x F_y + \frac{y\alpha}{m}F_z F_y - \frac{x\alpha}{m}F_z F_x + \frac{xy}{m}F_z^2 = 0, \\ \alpha F_{xz} - yF_{zz} + \alpha F_y + xF_z - \frac{\alpha}{m}F_x F_z + \frac{y}{m}F_z^2 = 0, \\ \alpha F_{yz} + xF_{zz} - \alpha F_x + yF_z - \frac{\alpha}{m}F_y F_z - \frac{x}{m}F_z^2 = 0, \end{cases}$$

where $\alpha = 1 + \frac{3x^2+3y^2}{4}$ and indices denote the derivative with respect to x , y and z .

Case A: Assume that the potential function F depends only on x , so (3.38) reduces to

$$(3.39) \quad \nu = 0, \quad F_x = 0.$$

Therefore, quasi-*Einstein soliton is steady and F is constant.

Case B: Now, assume that the potential function F depends only on y , so (3.38) reduces to

$$(3.40) \quad \nu = 0, \quad F_y = 0.$$

That is, quasi-*Einstein soliton is steady and F is constant.

Case C: Further, assume that the potential function F depends only on z , so (3.38) reduces to

$$(3.41) \quad \nu = 0, \quad F_z = 0.$$

Which implies quasi-*Einstein soliton is steady and F is constant.

4 Quasi-*Einstein ($\kappa < 1, \mu$) spaces

In this section, we study quasi-*Einstein ($\kappa < 1, \mu$)-spaces.

For a (κ, μ) -space [2]

$$(4.1) \quad h^2 = -(1 - \kappa)\phi^2,$$

where $\kappa \leq 1$.

Theorem 4.1. [3] *Let $(M^{2n+1}, \xi, \eta, \phi, g)$ be a non-Sasakian (κ, μ) -space ($\kappa < 1$). Then its Riemann curvature tensor R is given by*

$$(4.2) \quad \begin{aligned} g(R(Z, V)Y, W) = & (1 - \frac{\mu}{2})R_1 + R_2 + \left(\frac{1 - \frac{\mu}{2}}{1 - \kappa}\right)R_3 - \frac{\mu}{2}R_4 + \left(\frac{\kappa - \frac{\mu}{2}}{1 - \kappa}\right)R_5 \\ & + \mu g(\phi Z, V)g(\phi Y, W) + \eta(Z)\eta(W)R_6 - \eta(Z)\eta(Y)R_7 \\ & + \eta(V)\eta(Y)R_8 - \eta(V)\eta(W)R_9, \end{aligned}$$

where

$$\begin{aligned} R_1 &= g(V, Y)g(Z, W) - g(Z, Y)g(V, W), \\ R_2 &= g(V, Y)g(hZ, W) - g(Z, Y)g(hV, W) \\ &\quad - g(V, W)g(hZ, Y) + g(Z, W)g(hV, Y), \\ R_3 &= g(hV, Y)g(hZ, W) - g(hZ, Y)g(hV, W), \\ R_4 &= g(\phi V, Y)g(\phi Z, W) - g(\phi Z, Y)g(\phi V, W), \\ R_5 &= g(\phi hV, Y)g(\phi hZ, W) - g(\phi hV, W)g(\phi hZ, Y), \\ R_6 &= (-1 + \kappa + \frac{\mu}{2})g(V, Y) + (-1 + \mu)g(hV, Y), \\ R_7 &= (-1 + \kappa + \frac{\mu}{2})g(V, W) + (-1 + \mu)g(hV, W), \\ R_8 &= (-1 + \kappa + \frac{\mu}{2})g(Z, W) + (-1 + \mu)g(hZ, W), \\ R_9 &= (-1 + \kappa + \frac{\mu}{2})g(Z, Y) + (-1 + \mu)g(hZ, Y), \end{aligned}$$

$\forall Z, V, Y, W \in TM$.

Using (1.2) and (4.2) we find that

$$(4.3) \quad Q^*Z = (\kappa + n\mu)\phi^2Z,$$

and *-scalar curvature is given by

$$(4.4) \quad r^* = -2n(n\mu + \kappa),$$

which is constant.

Theorem 4.2. *Let M^{2n+1} be a $(\kappa < 1, \mu)$ -space satisfying (1.5). Then, M is a *-flat or locally isometric to $E^{n+1} \times S^n(4)$ or *Einstein.*

Proof. From (4.3), we have

$$(4.5) \quad Q^*\xi = 0.$$

Differentiating (4.5) along $Z \in TM$ and using (2.9), we have

$$(4.6) \quad (\nabla_Z Q^*)\xi = Q^*\phi Z + Q^*\phi hZ.$$

Taking inner product of (3.3) with ξ and using (4.5), we obtain

$$(4.7) \quad \begin{aligned} g(R(Z, V)\nabla F, \xi) &= g((\nabla_V Q^*)Z - (\nabla_Z Q^*)V, \xi) \\ &\quad - \frac{\nu}{m}\{(ZF)\eta(V) - (VF)\eta(Z)\}. \end{aligned}$$

Using (2.11) and (4.6) in (4.7) and then replacing Z by ϕZ and V by ϕV , we get

$$(4.8) \quad (Q^*\phi + \phi Q^*)Z - hQ^*\phi Z - \phi Q^*hZ = 0,$$

$\forall Z \in TM$. Further, using (2.1) and (4.3) in (4.8), we find that

$$(4.9) \quad \kappa + n\mu = 0.$$

Putting $V = \xi$ in (4.7), using (2.11), $\langle (\nabla_Z Q^*)\xi, \xi \rangle = 0$, $\langle (\nabla_\xi Q^*)Z, \xi \rangle = 0$, we get

$$(4.10) \quad h\nabla F = \sigma(\nabla F - (\xi F)\xi),$$

where $\sigma = \frac{\nu - \kappa m}{m\mu}$ is constant. Differentiating (4.10) along $Z \in TM$ and using (2.9), (3.5) and (4.10), we obtain

$$(4.11) \quad \begin{aligned} &(\nabla_Z h)\nabla F - hQ^*Z - \frac{\sigma(\xi F)}{m}(ZF)\xi + \nu hZ \\ &= \sigma[\nu Z - Q^*Z - (Z(\xi F))\xi + (\xi F)(\phi Z + \phi hZ)]. \end{aligned}$$

Using (4.5) in (3.5), we get

$$(4.12) \quad \xi(\xi F) - \frac{1}{m}(\xi F)^2 = \nu.$$

On the other hand, using (2.11), we get $l = \mu h - \kappa \phi^2$. Using this and (4.1) in (2.10), we find

$$(4.13) \quad \nabla_{\xi} h = -\mu \phi h.$$

Putting ξ in place of Z in (4.11) and using (4.12) and (4.13), we find

$$(4.14) \quad \mu h \nabla F = 0.$$

From (4.14), we have either $\mu = 0$ or $\mu \neq 0$.

Case A: Let $\mu = 0$. Then using this in (4.9), we obtain $\kappa = 0$. Thus $R(Z, V)\xi = 0$ and hence in dimension 3, M is $*$ -flat and in higher dimension M is locally isometric to $E^{n+1} \times S^n(4)$. [1]

Case B: In this case, we have $h \nabla F = 0$. Using $h^2 = (\kappa - 1)\phi^2$, we get

$$(4.15) \quad 0 = h^2 \nabla F = (\kappa - 1)\phi^2 \nabla F.$$

Since $\kappa < 1$, we have $\nabla F = (\xi F)\xi$. Hence from Lemma 3.2, we get F is constant. Consequently from (3.5), $Q^*Z = \nu Z$. Hence M is $*$ -Einstein. \square

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